

Probabilistic Modeling for Solar Energetic Particle Events

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Overview

- New data base of Solar Energetic Particle (SEP) episodes for protons and helium
- Short and Long mission approaches for probabilistic modeling
- Examples of the peak flux and episode-integrated fluence probabilistic models

Data base of SEP Episodes

- An episode is one or more events that overlap.
 - Flux doesn't return to background in between events
- Episode-Integrated Fluence Proton Data base
 - GME on IMP-8 and EPS on GOES
 - Normalized using isotropic periods of flux and Rodriguez et al. [2014]
 - Redistributed the GOES fluence in GME channels

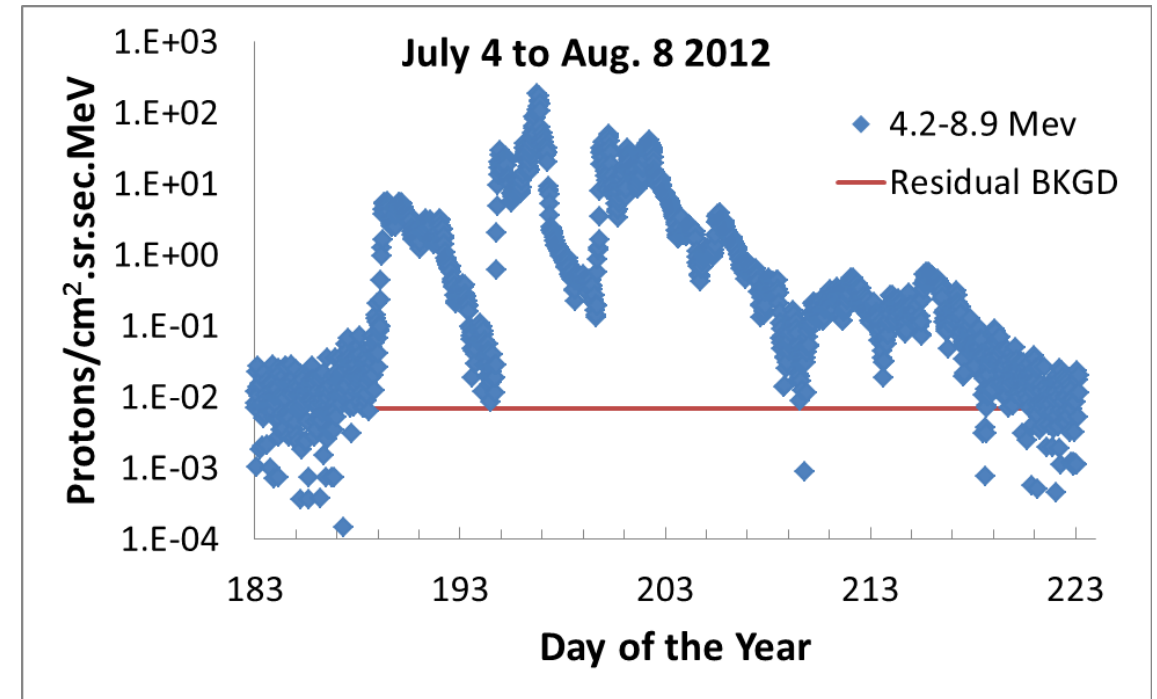
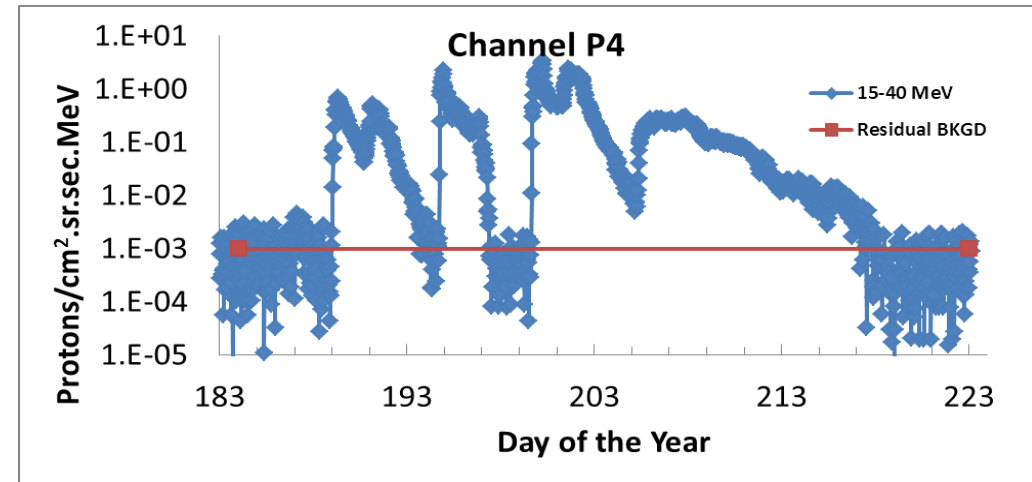
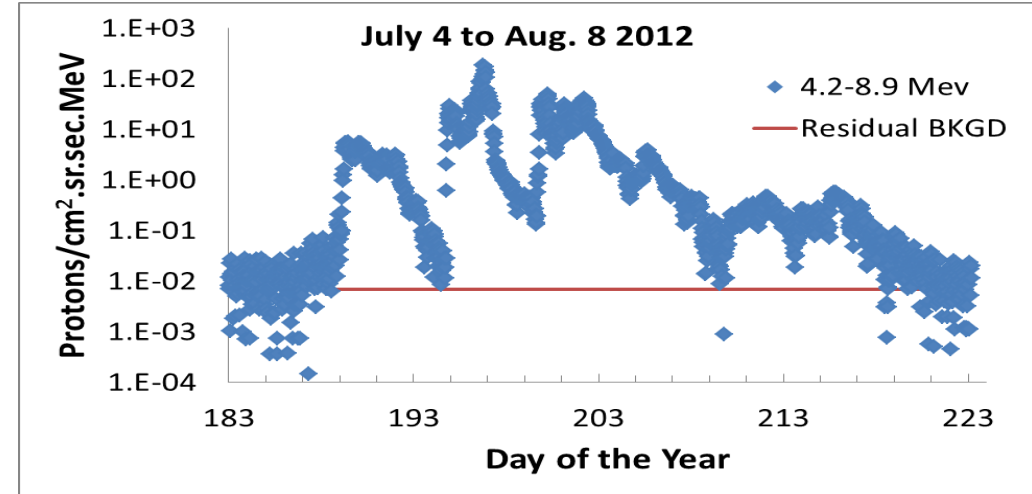


Image from
[Robinson, 2015]

Data Base of SEP Peak Fluxes

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- Peak Flux Data Base
 - Proton
 - EPS on GOES
 - Normalized using periods of isotropic flux and Rodriguez et al. [2014]
 - Helium
 - Solar Energetic Particle Environment Modeling (SEP-EM) system [Crosby et al., 2015]



Images from [Robinson, 2015]

Probabilistic Modeling

For long missions: [Xapsos et al., 1998]

The probability of no event with a flux $\geq \phi$ occurring in T years:

$$F_T(M) = \sum_n \frac{(\mu T)^n}{n!} \exp(-\mu T) [P(M)]^n$$

Where $M = \log(\phi)$

This can be simplified to :

$$F_T(M) = \exp\{-\mu T[1 - P(M)]\}$$

Probabilistic Modeling

Long Missions: [Xapsos et al., 1998]

$$F_T(M) = \exp\{-\mu T[1 - P(M)]\}$$

- Cumulative distributions built from data bases.
- Episodes per year
- User-specified mission start date, duration, and confidence level

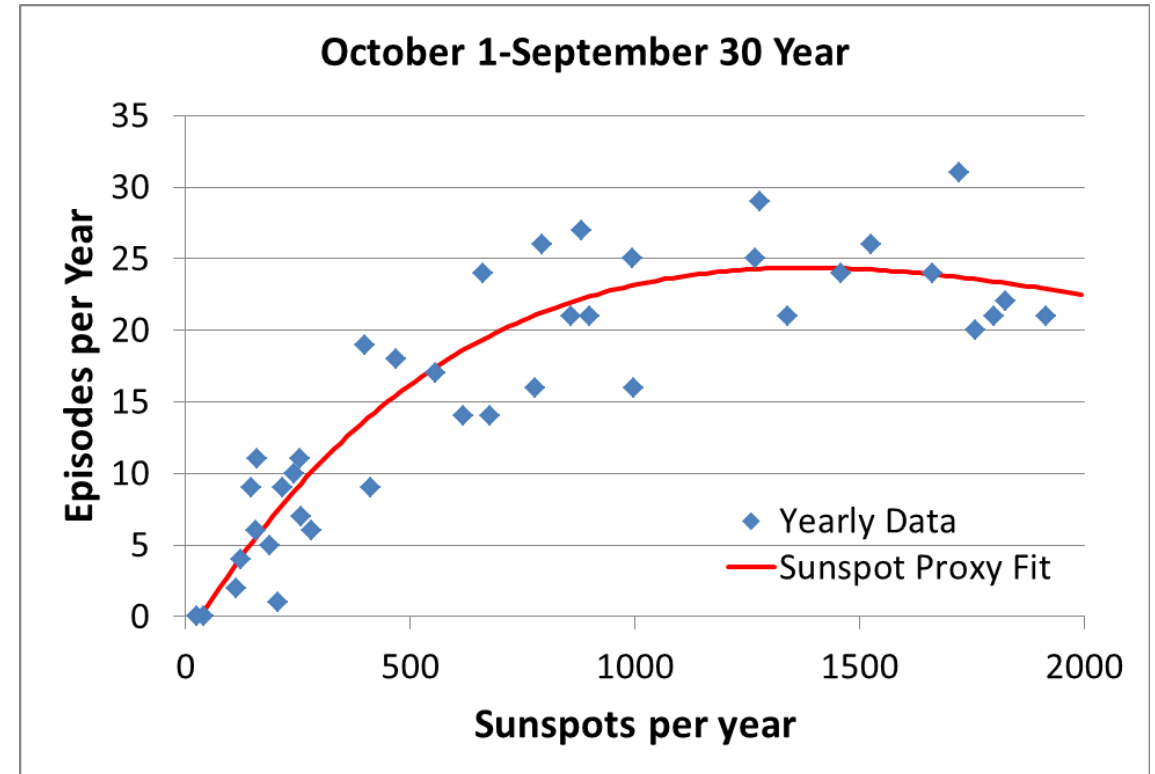
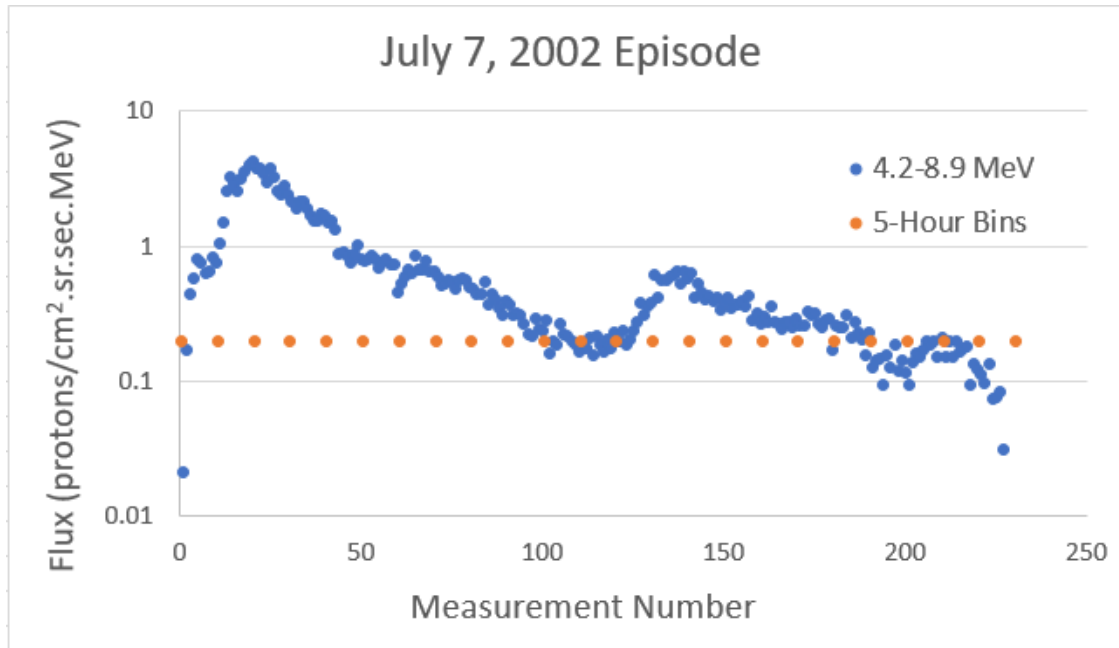


Image from
[Robinson, 2015]

Probabilistic Modeling

Short Mission:

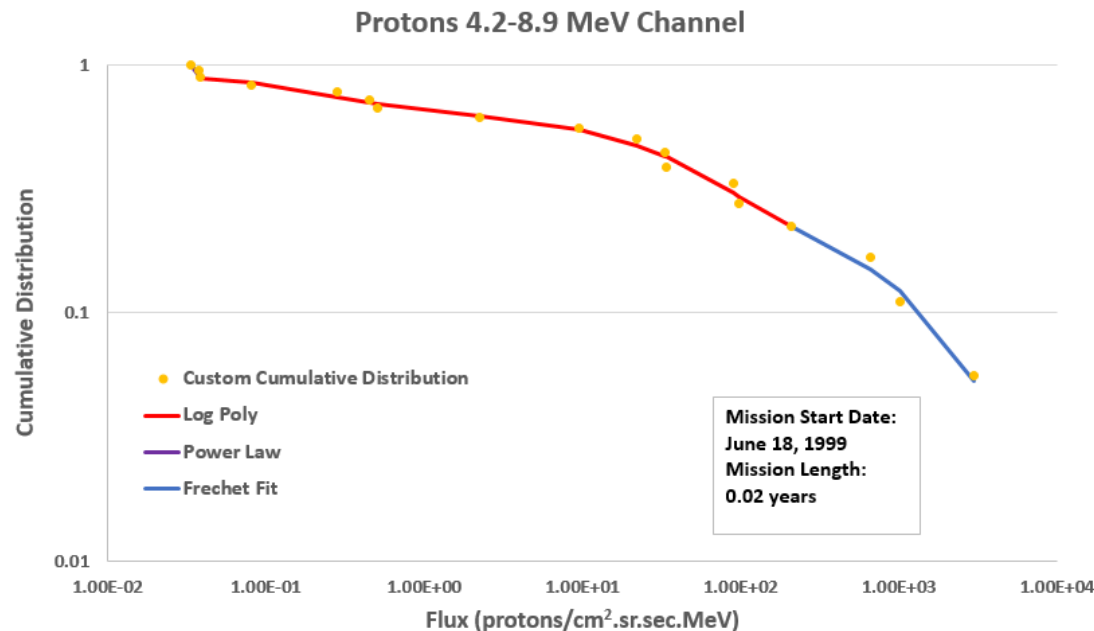
- Chronological list of flux measurements
 - Remove measurement at BKGD
- Mission length used to group data
 - Maximum flux taken to build a custom cumulative distribution
- Confidence Level used to determine if flux is above BKGD



Probabilistic Modeling

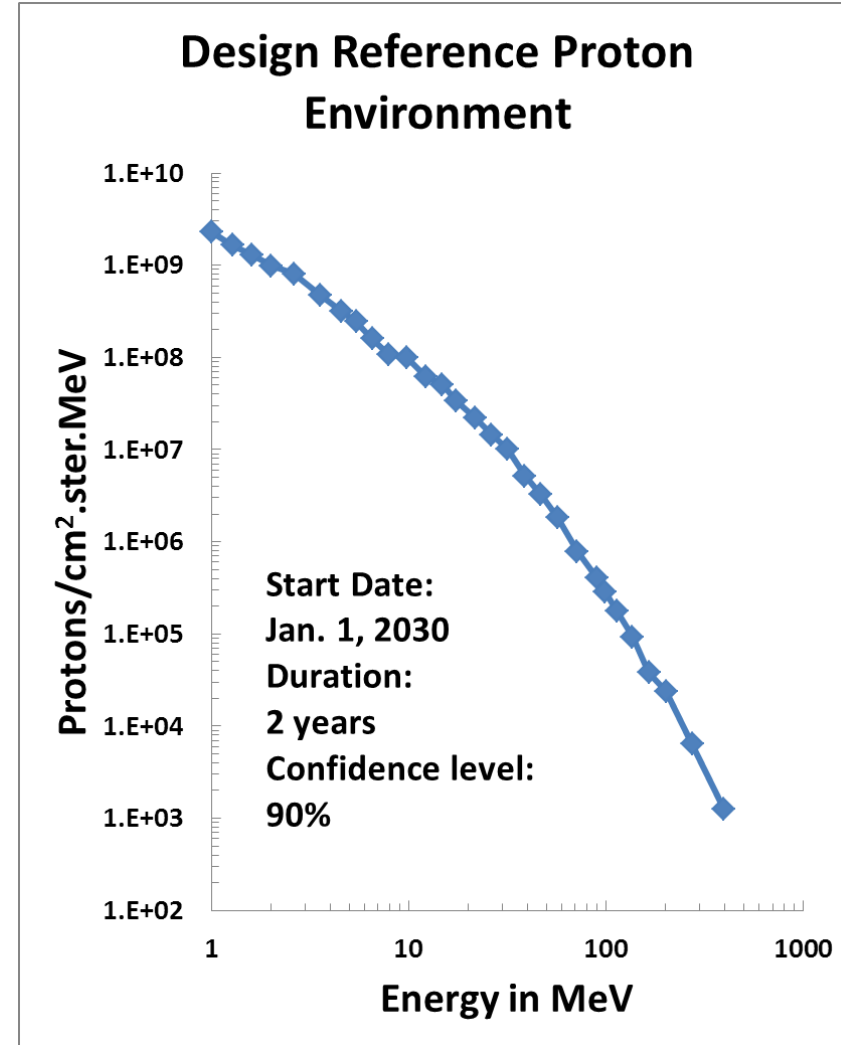
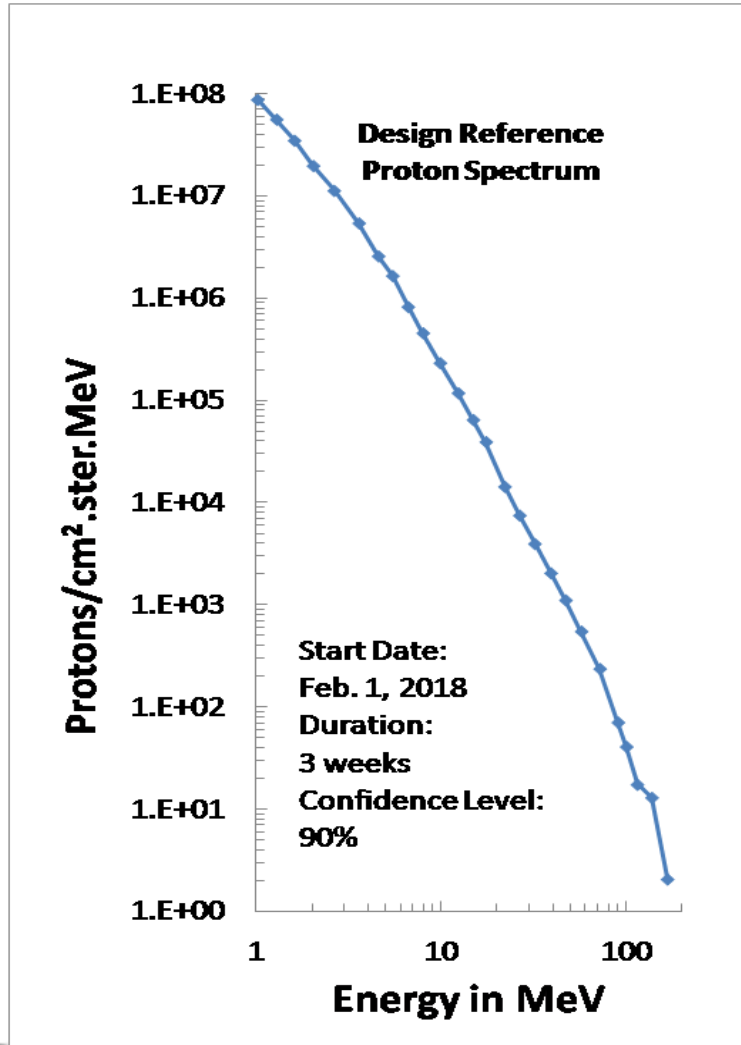
Short Mission:

- If under 1000 groups, use 3 fits [Robinson, 2015]:
 - Power Law
 - Logarithmic Polynomial
 - Fréchet Distribution
- If over 1000 groups, use the custom distribution
- Linear interpolation used between values in distribution



Episode-Integrated Fluence Model

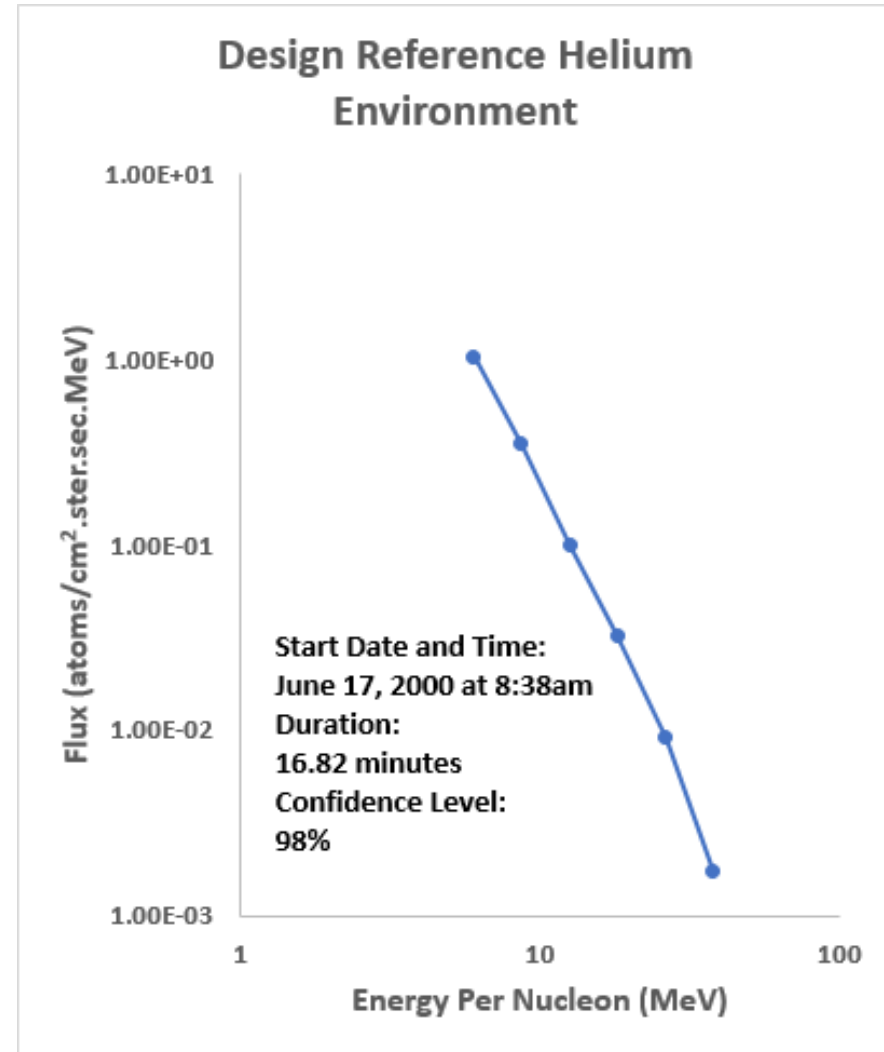
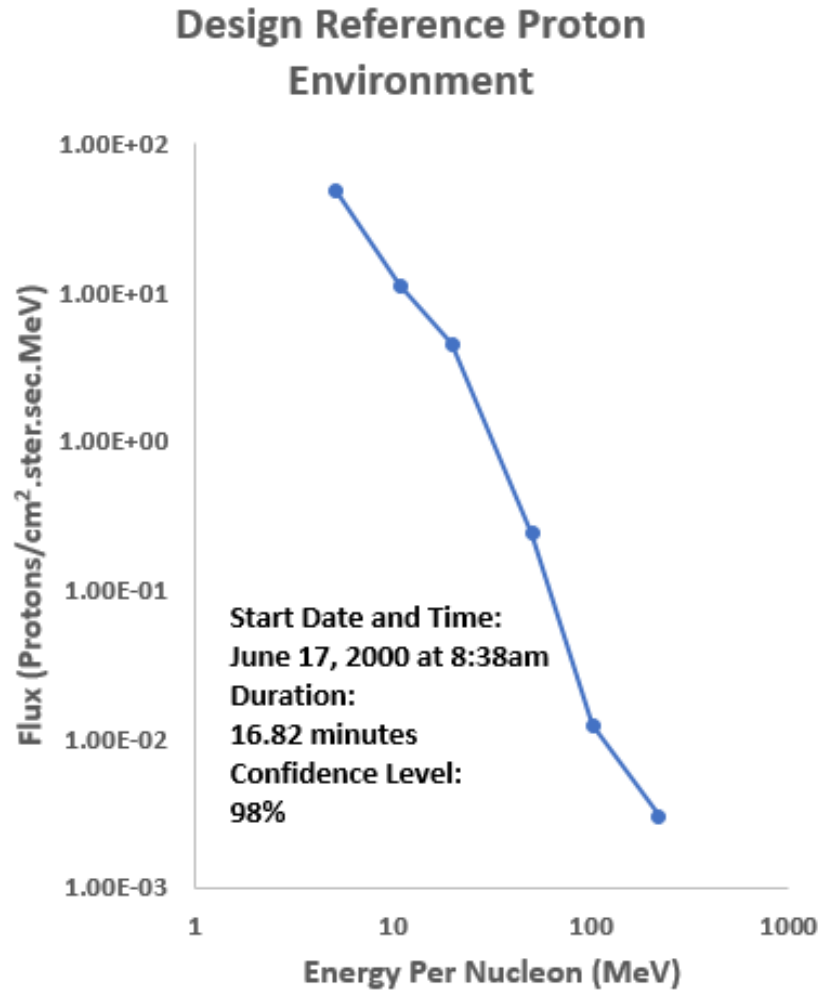
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Images from [Robinson, 2015]



Peak Flux Model

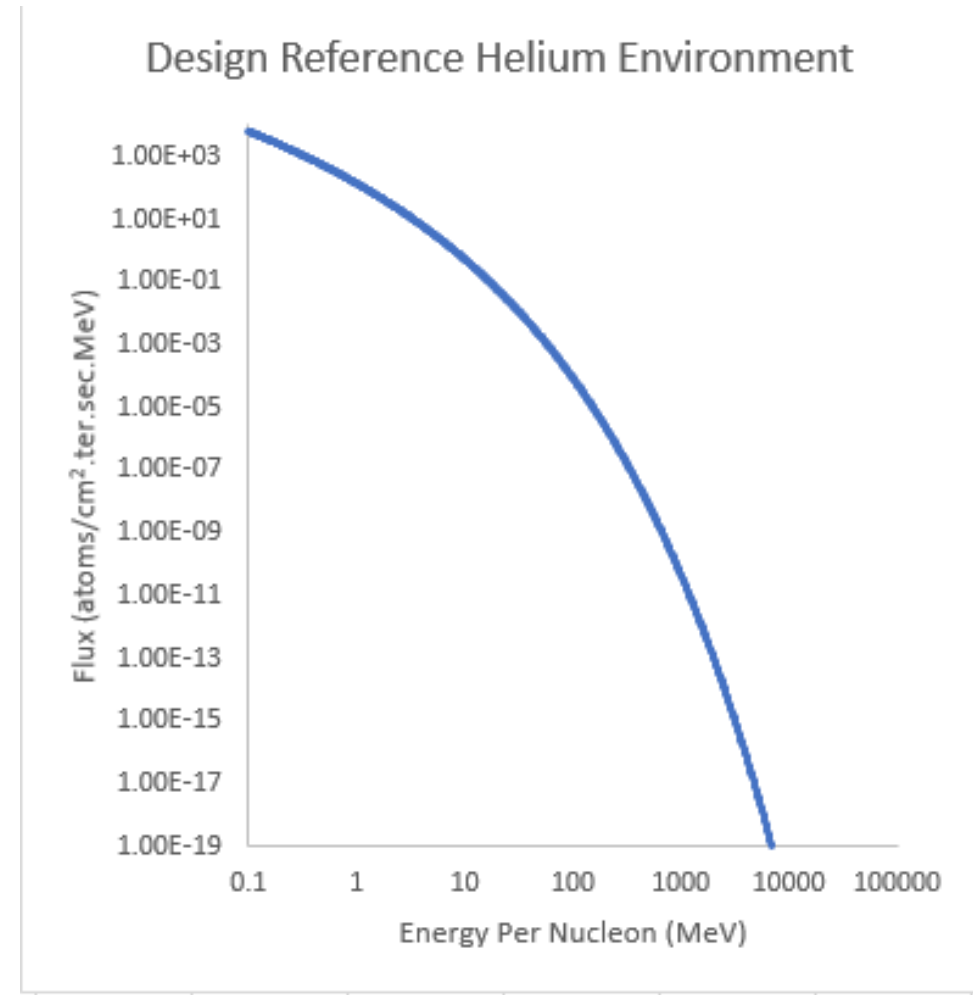
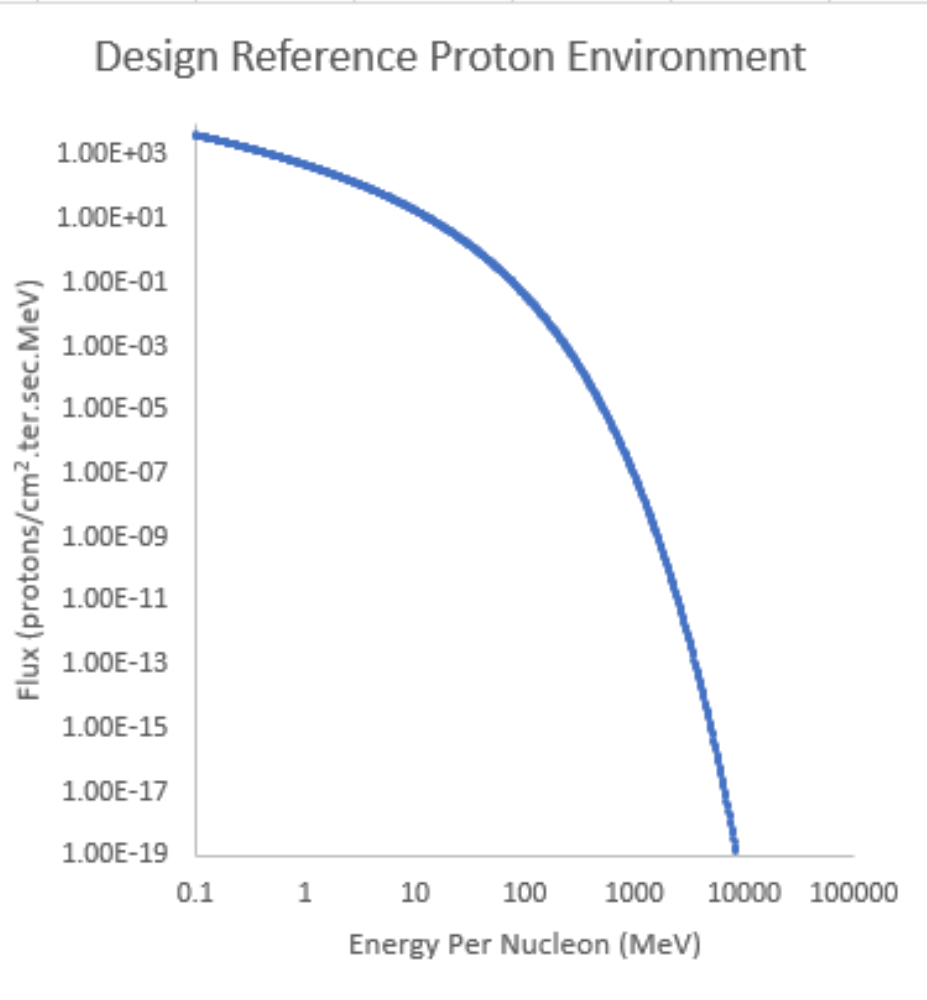


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Peak Flux Model in CREME96 Energy Channels

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Future Work

Episode-Integrated Fluence Model:

- Extend model to include 2014-2016 proton data.

Peak Flux Model:

- Create data bases for heavier elements and include Z=3-92 in model. Use measured elemental ratios to scale distributions.
- Integration into Space Ionizing Radiation Environments and Effects (SIRE²) toolkit [Fisher et al., 2017].
 - GUI based toolkit based on Cosmic Ray Effects on Micro-Electronics, 1996 version (CRÈME96)

References

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- Rodriguez, J. V., J. C. Krosschell, and J. C. Green, “Intercalibration of GOES 8-15 solar proton detectors,” *Space Weather*, Vol. 12, Iss. 1, 2012, 92–109.
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