

# Using Machine Learning to Detect Cloud Signatures in COSMIC-2 Radio Occultations

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## Objective

Explore the feasibility of using Machine Learning (ML) to detect clouds using Radio Occultation atmospheric profile data.

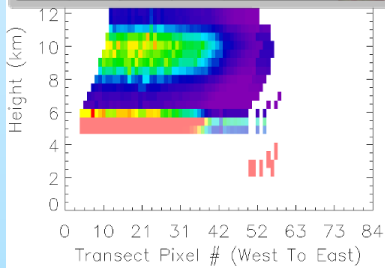
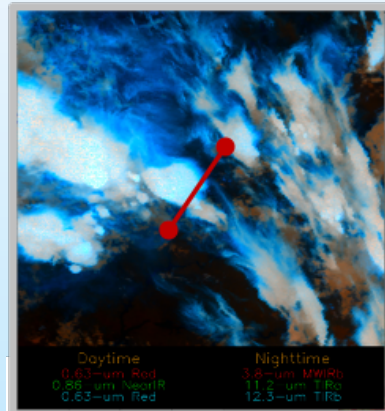
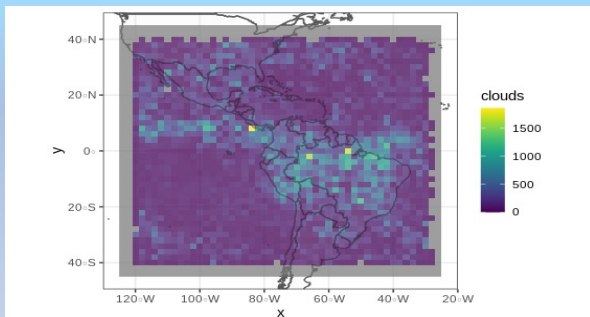
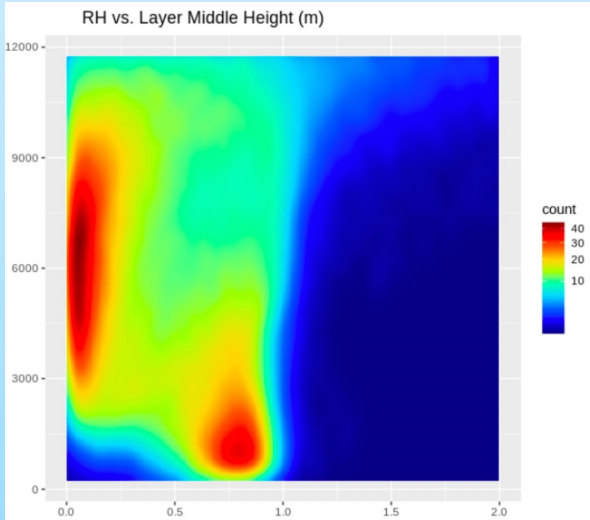
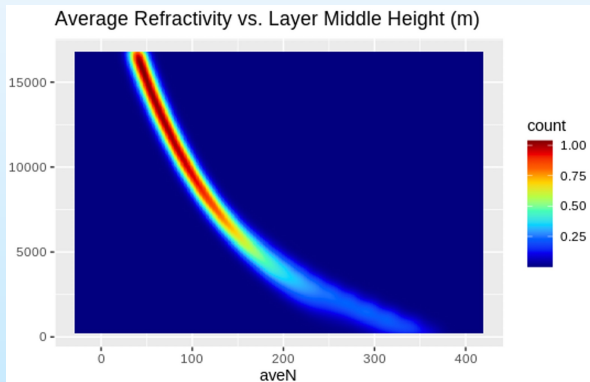
## Approach

The data used in this study comes from three primary sources:

### Inputs

1. Cosmic2 atmprf files
  1. Refractivity Profile
  2. Dry Temperature Profile
  3. Bending Angle Profile
  4. Altitude Profile
2. GFS NWP data
  1. Temperature Profile
  2. RH
3. GOES
  1. Use GOES channel data and NWP to map clouds to voxel locations
  2. Voxels are GOES pixels by 500m deep vertical sections from 6 to 12 km

Target for ML is cloud vs no cloud for voxels that the RO transect intersects



Confusion Matrix and Statistics

		Reference	
		0	1
Prediction	0	22738	2562
	1	2414	4776

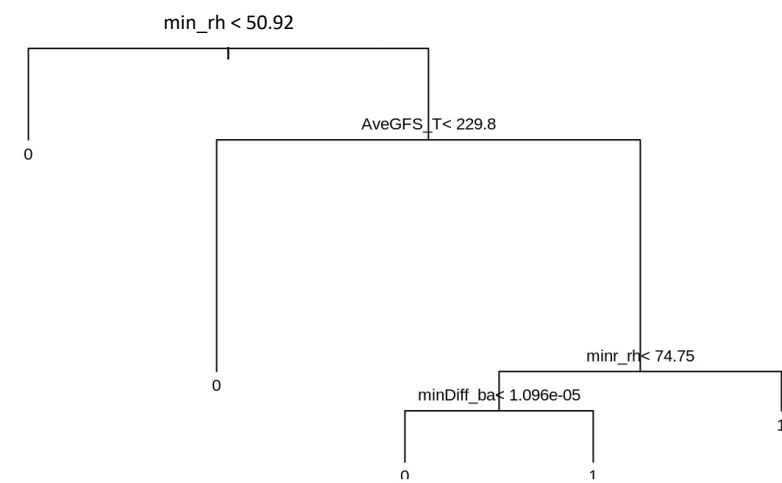
Accuracy : 0.8468  
 95% CI : (0.8429, 0.8507)  
 No Information Rate : 0.7741  
 P-Value [Acc > NIR] : < 2e-16

Kappa : 0.5589  
 McNemar's Test P-Value : 0.03717

Sensitivity : 0.9040  
 Specificity : 0.6509  
 Pos Pred Value : 0.8987  
 Neg Pred Value : 0.6643  
 Prevalence : 0.7741  
 Detection Rate : 0.6998  
 Detection Prevalence : 0.7787  
 Balanced Accuracy : 0.7774

'Positive' Class : 0

## Simple Decision Tree



## ROC and AUC for ML Models

