

Objective

To show how inner boundary condition (IBC) settings affect simulation results, which in turn, impact geo magnetic disturbances (GMD) forecasting.

Single Fluid vs. Multifluid MHD Equations

Single Fluid:

- Protons and electrons
- No mass separation of other ion species
- Single energy equation solved
- No collision allowed
- Less computational demand

Multifluid:

- Multiple ion species and electrons can be accounted for
- Separate energy equations for different ion species
- More computationally demanding for solving the ideal MHD multiple times

$$\frac{\partial \rho_m}{\partial t} + \nabla \cdot (\rho_m \mathbf{u}) = 0$$

$$\rho_m \frac{\partial \mathbf{u}}{\partial t} + \rho_m (\mathbf{u} \cdot \nabla) \mathbf{u} + \nabla p = \rho_m \mathbf{g} + \mathbf{j} \times \mathbf{B}$$

$$\frac{\partial \mathbf{B}}{\partial t} - \nabla \times (\mathbf{u} \times \mathbf{B}) = -\eta_m \nabla^2 \mathbf{B}$$

$$\frac{3}{2} \frac{\partial p}{\partial t} + \frac{3}{2} (\mathbf{u} \cdot \nabla) p + \frac{5}{2} p (\nabla \cdot \mathbf{u}) = \frac{1}{\sigma_0} \mathbf{j}^2$$

$$\frac{\partial p_s}{\partial t} + \nabla \cdot (\rho_s \mathbf{u}_s) = S_{p_s}$$

$$\frac{\partial \rho_s \mathbf{u}_s}{\partial t} + \nabla \cdot (\rho_s \mathbf{u}_s \mathbf{u}_s + I p_s) = n_s q_s (\mathbf{u}_s - \mathbf{u}_+) \times \mathbf{B} + \frac{n_s q_s}{n_e e} (\mathbf{J} \times \mathbf{B} - \nabla p_e) + S_{p_s}$$

$$\frac{\partial p_e}{\partial t} + \nabla \cdot (p_e \mathbf{u}_e) = -(\gamma - 1) p_e \nabla \cdot \mathbf{u}_e + S_{p_e}$$

$$\frac{\partial \mathbf{B}}{\partial t} - \nabla \times (\mathbf{u}_+ \times \mathbf{B}) = 0,$$

Toth et al., 2009

Importance: Allowing other ion species to exist in the MHD equations, this can impact reconnection rates, recovery rates post storm, and impact physics throughout the magnetosphere.

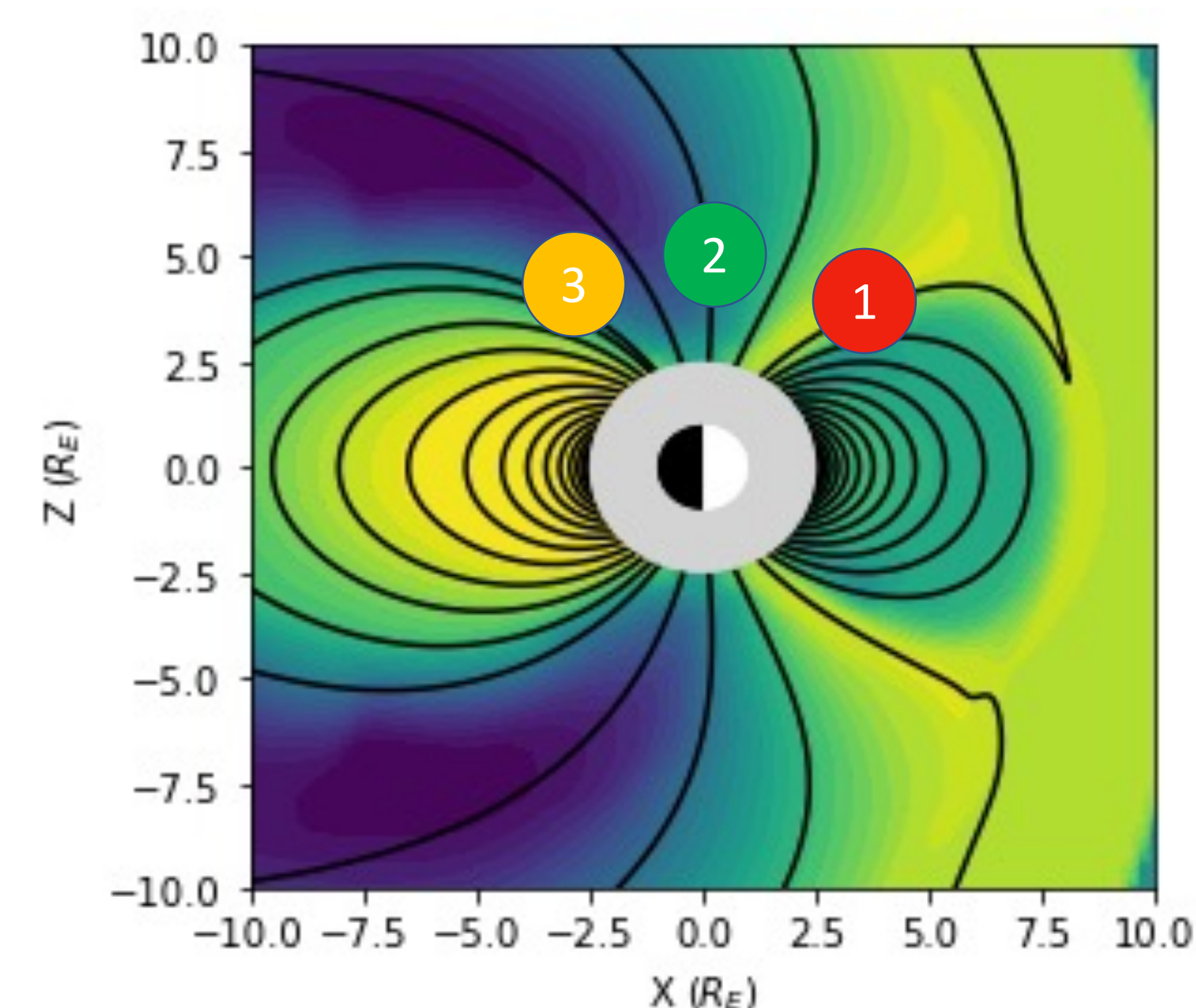
Methodology

Simulation set up:

+3 nT northward IMF B_z for 5 hours to -10 nT southward IMF B_z for 5 hours
Idealized run parameters, compared light, medium, heavy densities at inner boundary

- Coupled General Magnetosphere (GM) and Ionosphere Electrodynamics (IE) using the Space Weather Modeling Framework

Ion Outflow Regions of Interest



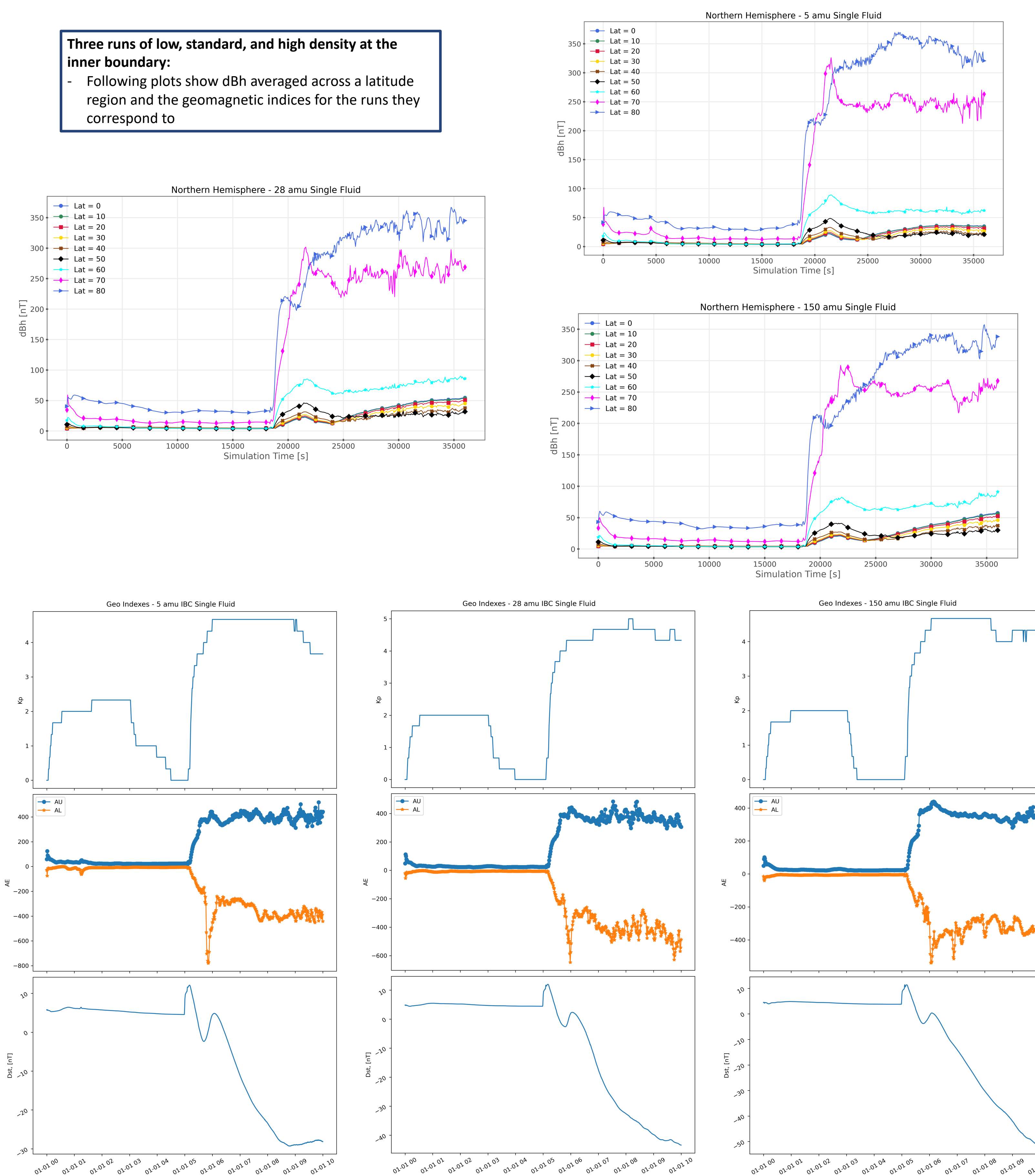
Regions of Interest:

1. Cusp Outflow
2. Auroral Outflow
3. Ring Current Outflow

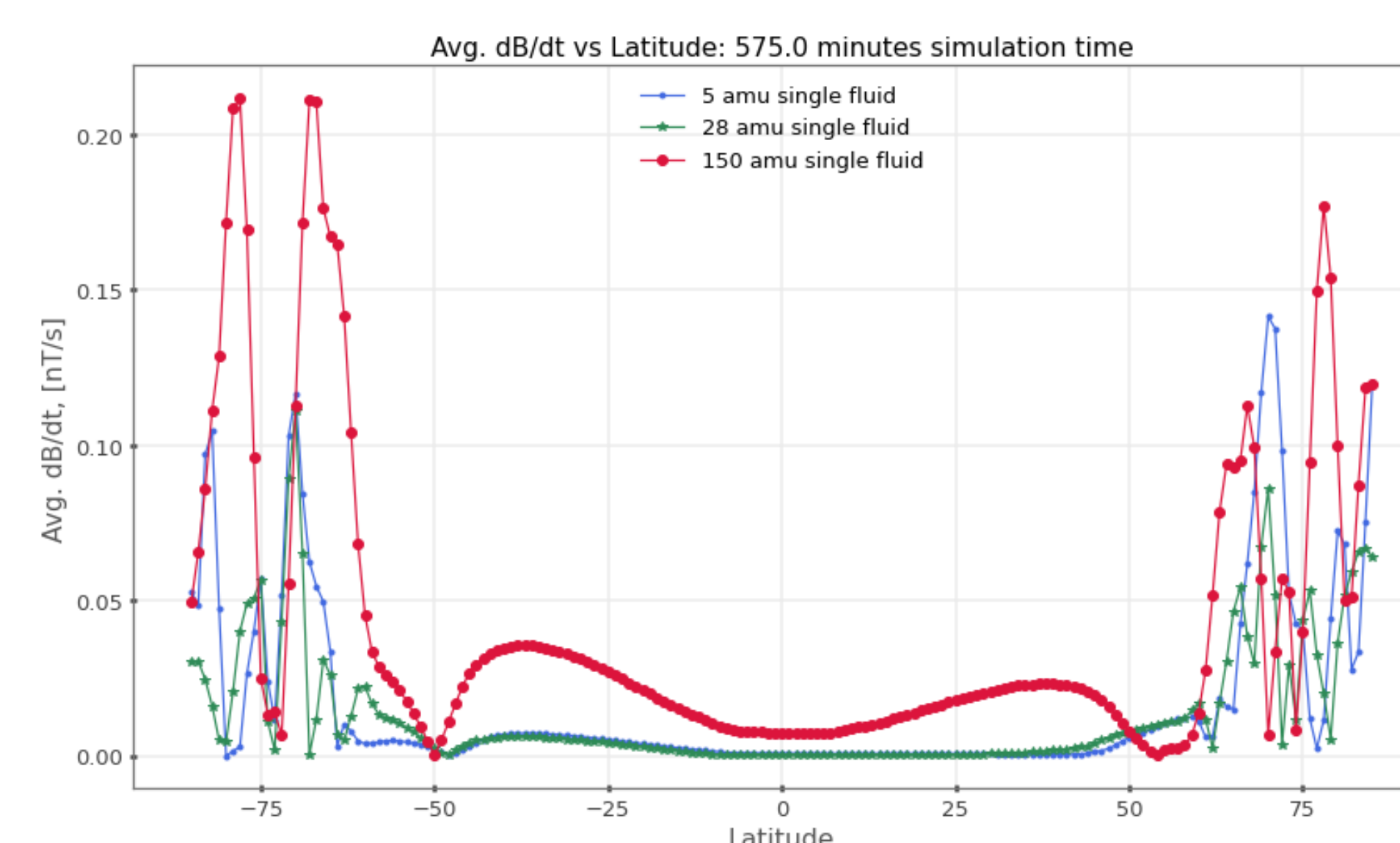
Results

Three Densities – Idealized Runs

Three runs of low, standard, and high density at the inner boundary:
 - Following plots show dBh averaged across a latitude region and the geomagnetic indices for the runs they correspond to



Averaged dB/dt for different densities



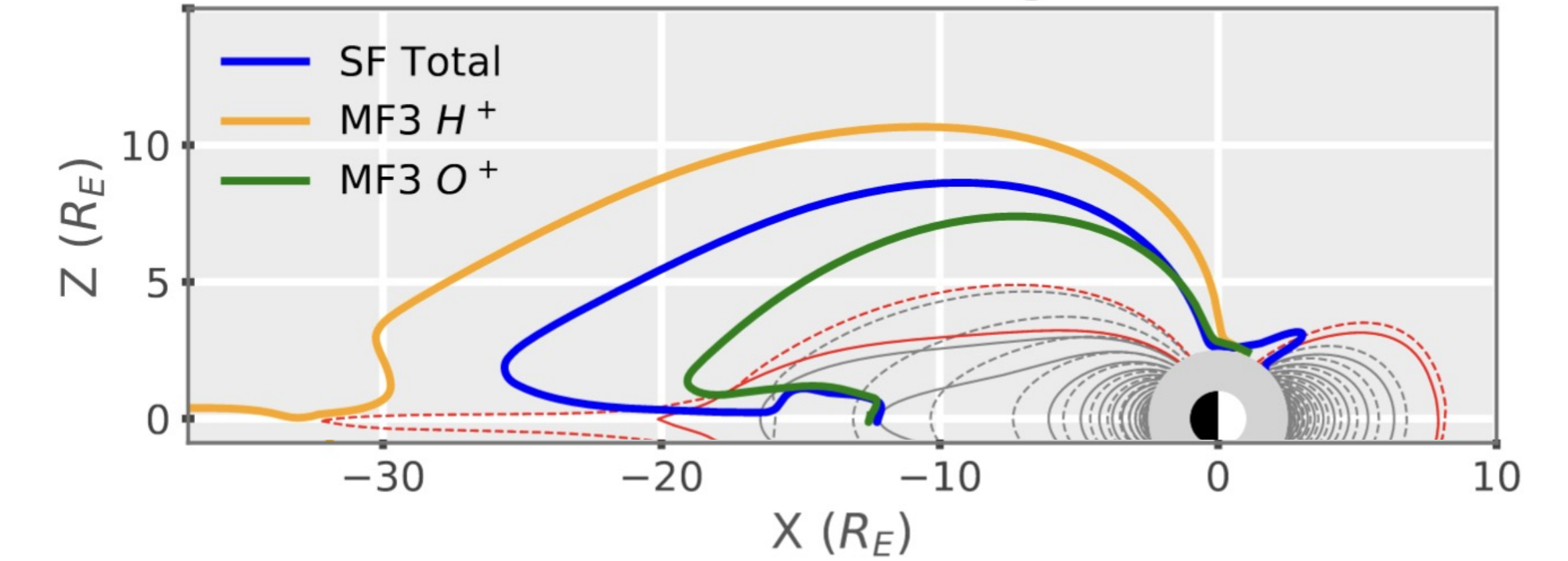
An increase in density leads to a further latitudinal extent!

Multifluid Impact

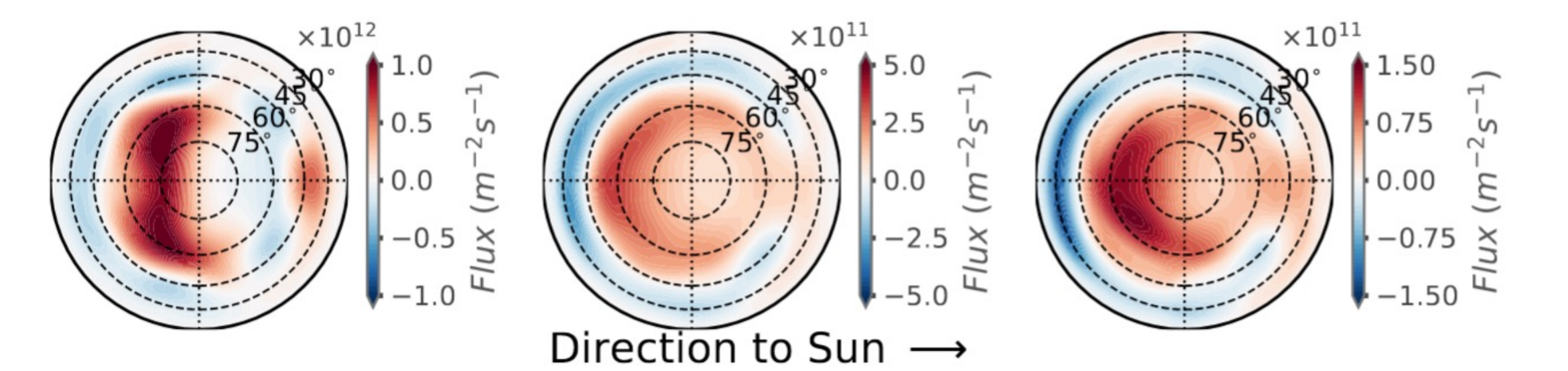
Multifluid:

- Multifluid run (shown below) indicates that with the presence of O⁺, this will feed into the ring current, while H⁺ will exit tailward.

Lobe Flow: MF3 vs. Single Fluid



The presence of O⁺ will change the outflow pattern which dictates the ion trajectory throughout the magnetosphere



Conclusions

Increasing the inner boundary density impacts the latitudinal extent of the average dB/dt. This, in turn, affects GMD forecasting.

It is anticipated that multifluid MHD would further change these results by

Next Steps

Fully flush out multifluid extent of the impact on GMD operations

- Include other IBCs like parallel velocity, temperature, or different compositions on top of number density change.

Compare results to a simulated magnetometer station for a case study

Compare H⁺/O⁺ ratios during storm time to determine accuracy of single fluid to multifluid MHD for GMD forecasting operations improvement.

Contact Information

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